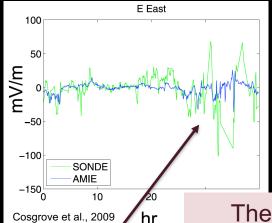


<u>Dogacan S. Ozturk (JPL)</u>, Xing Meng (JPL), Olga Verkhoglyadova (JPL), Michael Hartinger (SSI), Josh Semeter (BU), Roger Varney (SRI), Ashton Reimer (SRI)

contact: dogacan.s.ozturk@jpl.nasa.gov

M-I-T Coupling on Meso-Scale

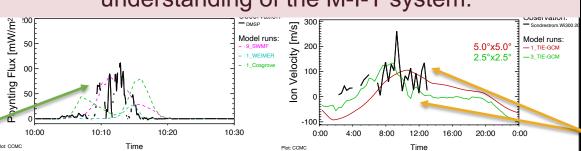


Assimilative models depend on coverage and resolution of the measurements

Both MHD and empirical models miss the observed variability (~1 min.) Global Circulation Models (GCMs) traditionally use empirical models for global estimates of electric fields and conductivity and significant work is ongoing to resolve mesoscale structures¹.

Missing meso-scale electric field variability (temporal + spatial)
causes underestimation of energy input and dissipation in
the high-latitude lonosphere².

There is a need to incorporate dynamic driving through meso-scale structures to further our understanding of the M-I-T system.



Model results improve on a finer grid, still underestimating the values.

t al. (2009); Cousins et al. (2013)

al. (2014): Brinkman et al. (2016)

Rastätter, et al., 2016

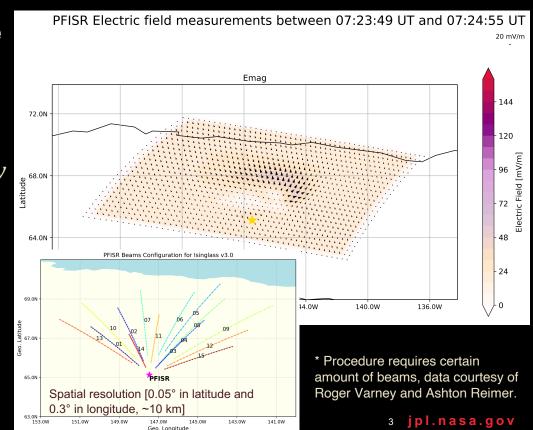
Incorporating meso-scale variability in a GCM

PFISR LOS velocity measurements can be used to derive Electric fields on a 2D grid*.

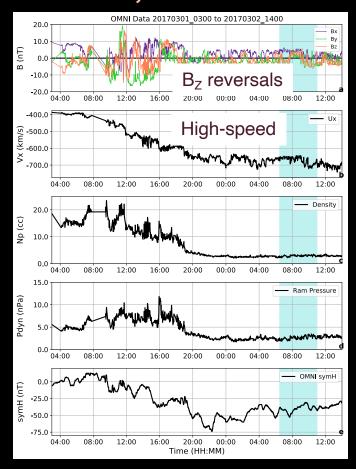
→ The potential change in longitude (x) and latitude (y) can be calculated:

$$\Delta\phi_x = -\int_{x_1}^{x_2} E_x \, dx \, , \, \Delta\phi_y = -\int_{y_1}^{y_2} E_y \, dy$$

- 1. Down-sample the x and y components of the electric fields on desired grid. (0.75°x0.75°)
- 2. Calculate the potentials on the new grid to drive GITM.

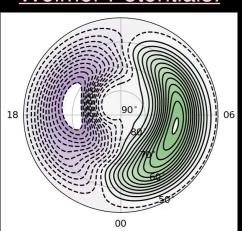


March 2, 2017: Solar wind and IMF Conditions

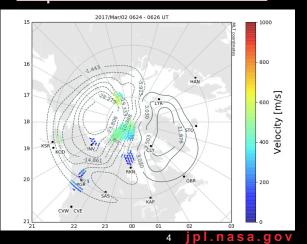


- PFISR was aiding Ionospheric Structuring: In Situ and Groundbased Low Altitude Studies (ISINGLASS) Experiment [Clayton et al., 2019]
- Recovery phase of a substorm
- No significant response in Weimer or SuperDARN potentials around 0630 UT.

Weimer Potentials:

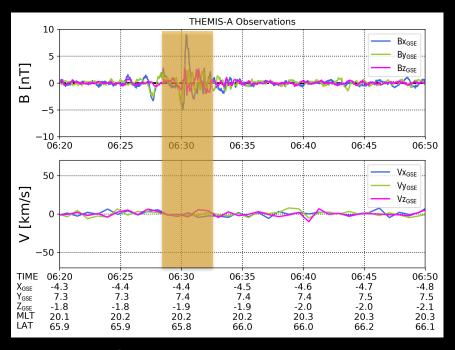


SuperDARN Potentials:

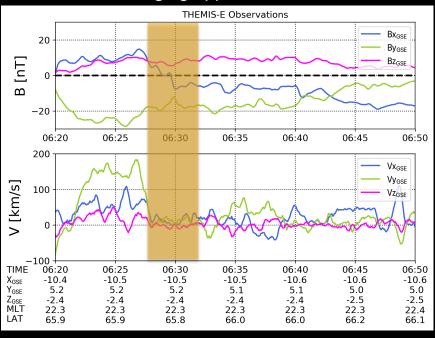


6/25/19

Magnetotail Activity

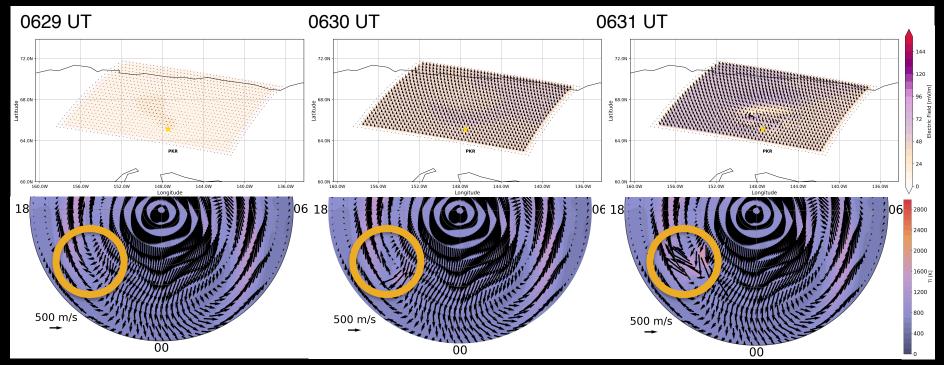


5-min. averaging applied to B and V values.



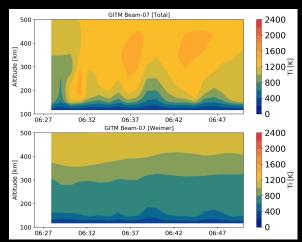
- THEMIS-A B_x measurements show strong fluctuations around 0630 UT.
- Around 0630 UT, THEMIS-E measurements show a B_X reversal.
- Measurements indicate localized ionospheric responses are possible.

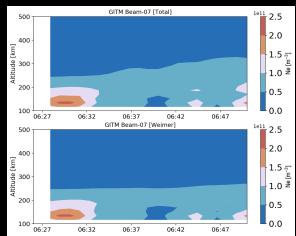
PFISR Electric Field Measurements – GITM Ion Temperature and Convection Simulations

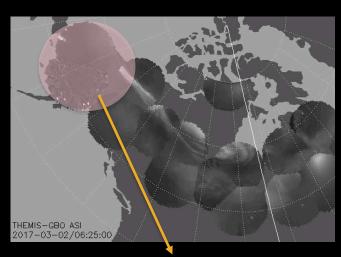


- Ion convection patterns inside PFISR region are enhanced in the westward direction.
- Convection further perturbed at 0631 UT, leading to an ion temperature enhancement of \sim 1000K.

Electron Density Response







THEMIS ASI images show impulsive brightening around 0630 UT

- Ion temperature response is immediate, penetrates to low altitudes.
- Electron density response between Weimer and meso-scale driven run are inseparable, however responses differ later in time.
- Lack of meso-scale particle precipitation in the model → lack of electron density response.
- Working on validation with PFISR (low SNR) and GPS measurements.

Summary

Conclusions:

- We are developing a framework that can utilize high-latitude local (meso-scale) 2D electric field measurements as input to run a global I-T model.
- Meso-scale structures are responsible for enhancements in ion temperature and perturbations in ion convection profiles.

Future work:

- Investigate the effects of meso-scale electric fields on the global energy budget during active geomagnetic periods.
- Include a self-consistent treatment of particle precipitation and electrodynamics for a complete understanding of meso-scale variability.
- Validation Studies: More events, more conjunctions, different sets of drivers and measurements to validate results
- Error and uncertainty quantification

Thank you.

Acknowledgements

- This work is funded by the NASA ROSES 2016 Heliophysics LWS Science (NRA NNH16ZDA001N) Program.
- GITM is developed and supported by Prof. Aaron Ridley at University of Michigan.
- Simulations were done on NASA High-End Computing Program through the NASA Advanced Supercomputing Division at Ames Research Center, and Stampede Supercomputer at Texas Advanced Computing Center at University of Texas at Austin.
- We acknowledge use of NASA/GSFC's Space Physics Data Facility's OMNIWeb (or CDAWeb or ftp) service, and OMNI data.
- We acknowledge NASA contract NAS5-02099 and V. Angelopoulos for use of data from the THEMIS Mission. Specifically: C. W. Carlson and J. P. McFadden for use of ESA data; K. H. Glassmeier, U. Auster and W. Baumjohann for the use of FGM data provided under the lead of the Technical University of Braunschweig and with financial support through the German Ministry for Economy and Technology and the German Center for Aviation and Space (DLR) under contract 50 OC 0302.
- We acknowldge the use of the VT SuperDARN Web site (vt.superdarn.org) for access to radar data and data products. SuperDARN is a collection of radars funded by national scientific funding agencies of Australia, Canada, China, France, Italy, Japan, Norway, South Africa, United Kingdom and the United States of America.
- Simulation results have been provided by the Community Coordinated Modeling Center at Goddard Space Flight Center through their public Runs on Request system (http://ccmc.gsfc.nasa.gov). The Weimer Model was developed by Daniel R. Weimer at Virginia Tech. The Ovation Prime Model was developed by Patrick Newell at JHU/APL.



jpl.nasa.gov